

# Matter-Wave Interferometry Tests of Density Field Dynamics

Gary Alcock<sup>1</sup>

<sup>1</sup>*Independent Researcher, Los Angeles, CA, USA*

(Dated: September 18, 2025)

Density Field Dynamics (DFD) posits a scalar refractive field  $\psi(\mathbf{x})$  such that light propagates with  $n = e^\psi$  (ONE-WAY phase speed  $c_1 = ce^{-\psi}$ ) and matter accelerates as  $\mathbf{a} = \frac{c^2}{2}\nabla\psi$ . While our cavity–atom redshift test probes the photon sector, *matter-wave interferometers* test the external wavefunction coupling. We derive the perturbative phase from the  $\nabla\psi\cdot\nabla$  operator in the DFD-modified Schrödinger equation and obtain a clean discriminator for light-pulse interferometers:

$$\Delta\phi_{\text{DFD}} = \frac{\hbar k_{\text{eff}}^2}{m} \frac{g}{c^2} T^3,$$

in contrast to the standard GR scaling  $\Delta\phi_{\text{GR}} = k_{\text{eff}}gT^2$ . We provide explicit, plug-in predictions for Kasevich–Chu, Raman, and Bragg geometries (vertical and horizontal), source-mass configurations, and dual-species protocols (Rb/Yb), and we analyze systematics with look-alike time scalings. For Earth  $g$  and  $k_{\text{eff}} \sim 1.6 \times 10^7 \text{ m}^{-1}$  (Rb, 780 nm), the DFD residual is  $\sim 2 \times 10^{-11}$  rad at  $T = 1$  s, within the reach of current long-baseline instruments when using rotation,  $k$ -reversal, and source-mass modulation.

## I. INTRODUCTION

Atom interferometers are leading probes of gravity, redshift, and fundamental symmetries.[1–6] In DFD, photons follow the eikonal of an optical metric with  $n = e^\psi$  while matter sees the conservative potential  $\Phi = -\frac{c^2}{2}\psi$ .<sup>1</sup> The photon-sector discriminator is a co-located cavity–atom redshift comparison across altitude; here we develop the *matter-sector* analogue: light-pulse atom interferometry. The novelty is a gradient–gradient coupling that yields a  $T^3$  scaling distinct from the GR  $T^2$  law, giving a route to sector-resolved falsification with existing facilities.

*Relation to existing gravity-gradient cancellation and why it was not seen.* Long-baseline experiments *actively suppress or calibrate out* cubic-in- $T$  gravity-gradient contributions using frequency-shift gravity-gradient (FSGG) compensation or closely related  $k$ -vector tuning schemes,[27–30] because within GR such terms are treated as *systematics*. As a result, published analyses typically (i) operate at fixed  $T$  for the headline measurement, (ii) do not report a residual *vs.*  $T$  regression with the *even-in- $k_{\text{eff}}$* , rotation-odd discriminator posed here, and (iii) use  $k$ -reversal specifically to cancel odd-in- $k_{\text{eff}}$  laser/systematic terms. To our knowledge, no experiment has isolated a coefficient  $b_{\text{even}}$  in  $\phi_{\text{res}}(T) = aT^2 + b_{\text{even}}T^3$  that (a) is *even* under  $k_{\text{eff}} \rightarrow -k_{\text{eff}}$  and (b) *flips sign* under 180° rotation of a horizontal baseline—the specific signature predicted here.

## II. THEORY: $\psi$ -COUPLING IN THE SCHRÖDINGER DYNAMICS

To first order in weak fields ( $|\psi| \ll 1$ ), the nonrelativistic equation for mass  $m$  reads (expanding  $e^{-\psi} \approx 1 - \psi$ )

$$i\hbar\partial_t\Psi = -\frac{\hbar^2}{2m}\nabla^2\Psi + m\Phi\Psi + \frac{\hbar^2}{2m}\left[\psi\nabla^2\Psi + (\nabla\psi)\cdot\nabla\Psi\right], \quad (1)$$

with  $\Phi \equiv -\frac{c^2}{2}\psi$ . Treat  $H = H_0 + \delta H$  with  $H_0 = \frac{p^2}{2m} + m\Phi$  and

$$\delta H = \frac{\hbar^2}{2m}\left[\psi\nabla^2 + (\nabla\psi)\cdot\nabla\right]. \quad (2)$$

Evaluate the small phase along the unperturbed classical branches  $A, B$ :

$$\Delta\phi_{\text{DFD}} = \frac{1}{\hbar} \int_0^{2T} dt \left( \langle\delta H\rangle_A - \langle\delta H\rangle_B \right). \quad (3)$$

The operator  $(\nabla\psi)\cdot\nabla$  acting on a locally plane-wave factor on each branch pulls down the instantaneous momentum,  $\langle(\nabla\psi)\cdot\nabla\rangle \rightarrow i(\nabla\psi)\cdot\mathbf{p}/\hbar$ , so that

$$\Delta\phi_{\nabla\psi} = -\frac{1}{2m} \int_0^{2T} dt (\nabla\psi)\cdot\Delta\mathbf{p}(t). \quad (4)$$

In uniform Earth gravity,  $\nabla\psi = -2\mathbf{g}/c^2$ ; the constant part cancels between arms unless one accounts for the *finite spatial separation* of the arms induced by the light pulses. Keeping the leading variation sampled at the arm positions yields the  $T^3$  law below.

<sup>1</sup> See the Einstein 1911–12 completion and the strong-field/GW manuscripts for the action, normalization, and recovery of GR’s weak-field coefficients; we adopt that notation here.

### III. LIGHT-PULSE GEOMETRIES AND THE $T^3$ DISCRIMINATOR

Consider a vertical Kasevich–Chu sequence ( $\pi/2$ – $\pi$ – $\pi/2$  at  $t = \{0, T, 2T\}$ ) with effective Raman wavevector  $k_{\text{eff}}\hat{z}$ . Let the recoil velocity be  $v_r = \hbar k_{\text{eff}}/m$ . Between pulses, the branch momentum difference is piecewise constant:  $\Delta p_z(t) = +\hbar k_{\text{eff}}$  for  $0 < t < T$ , and  $-\hbar k_{\text{eff}}$  for  $T < t < 2T$  (mirror swaps the arms). Using (4) with  $\nabla\psi(\mathbf{r}_{A,B}, t) = -2g\hat{z}/c^2$  evaluated at the *arm locations* and expanding to first order in the instantaneous arm separation  $\Delta z(t)$  (which is  $v_r t$  on the first half and  $v_r(2T-t)$  on the second), the constant part cancels but the linear piece adds over the two intervals, giving

$$\Delta\phi_{\text{DFD}}^{\text{Kasevich--Chu}} = \frac{k_{\text{eff}} v_r g}{c^2} T^3 = \frac{\hbar k_{\text{eff}}^2}{m} \frac{g}{c^2} T^3. \quad (5)$$

By contrast, the standard light-pulse phase from GR (after the usual laser phase bookkeeping) is

$$\Delta\phi_{\text{GR}}^{\text{Kasevich--Chu}} = k_{\text{eff}} g T^2. \quad (6)$$

*Numerics (Rb, 780 nm):*  $k_{\text{eff}} \simeq 1.6 \times 10^7 \text{ m}^{-1}$ ,  $v_r = \hbar k_{\text{eff}}/m \simeq 1.2 \times 10^{-2} \text{ m s}^{-1}$ . For  $T = 1 \text{ s}$ ,

$$\Delta\phi_{\text{DFD}}^{\text{Kasevich--Chu}} \approx \frac{(1.6 \times 10^7)(1.2 \times 10^{-2})(9.8)}{(3.0 \times 10^8)^2} \simeq 2 \times 10^{-11} \text{ rad}. \quad (7)$$

The absolute GR phase  $k_{\text{eff}} g T^2 \sim 1.6 \times 10^8 \text{ rad}$  is removed by chirp/common-mode subtraction; the *residual* DFD term is what to search for, using scaling and sign tests below.

#### A. Horizontal baselines and rotation

For a horizontal Raman/Bragg baseline with separation direction  $\hat{\mathbf{n}}$ , Earth’s field projects as  $\mathbf{g} \cdot \hat{\mathbf{n}}$ :

$$\Delta\phi_{\text{DFD}}^{\text{horiz}} = \frac{\hbar k_{\text{eff}}^2}{m} \frac{\mathbf{g} \cdot \hat{\mathbf{n}}}{c^2} T^3, \quad (8)$$

which *flips sign* under  $180^\circ$  rotation about the vertical. This provides a powerful discriminator from many systematics.

#### B. Source-mass configuration (tabletop)

Place a dense source mass (e.g.  $\sim 500 \text{ kg W}$ ) at distance  $R$  producing  $g_s = GM/R^2$ . Then

$$\Delta\phi_{\text{DFD}}^{\text{src}} = \frac{\hbar k_{\text{eff}}^2}{m} \frac{g_s}{c^2} T^3 \times \mathcal{G}(\text{geometry}), \quad (9)$$

where  $\mathcal{G}$  encodes near-field placement; lock-in by modulating the mass.

### C. Dual-species protocol (Rb/Yb)

Because the DFD term scales as  $\Delta\phi_{\text{DFD}} = (\hbar k_{\text{eff}}^2/m)(g/c^2)T^3$ , the differential phase between two species  $i, j$  operated in matched geometry is

$$\Delta\phi_{\text{DFD}}^{(i-j)} = \frac{gT^3}{c^2} \hbar \left( \frac{k_{\text{eff},i}^2}{m_i} - \frac{k_{\text{eff},j}^2}{m_j} \right). \quad (10)$$

If both species share the *same* lattice/Bragg wavelength (engineered co-propagating optics),  $k_{\text{eff},i} = k_{\text{eff},j}$  and (10) reduces to a clean mass discriminator  $\propto (1/m_i - 1/m_j)$ . With independent Raman pairs (e.g.  $^{87}\text{Rb}$  at 780 nm and  $^{171}\text{Yb}$  at 556 nm), keep the explicit  $k_{\text{eff}}$  values; Eq. (10) is then the quantity to regress against  $T^3$ . In either case, the GR common-mode  $k_{\text{eff}} g T^2$  cancels under standard  $k$ -reversal and conjugate-AI subtraction.

### IV. CONCRETE EXPERIMENTAL DESIGNS (PLUG-AND-PLAY)

*Design A (vertical Kasevich–Chu, 10 m fountain).* Species  $^{87}\text{Rb}$ ,  $\lambda = 780 \text{ nm}$ ,  $k_{\text{eff}} \approx 1.6 \times 10^7 \text{ m}^{-1}$ , pulses at  $t = \{0, T, 2T\}$  with  $T = 1\text{--}2 \text{ s}$ . Arm apex separation  $\Delta z_{\text{max}} \approx v_r T \sim 1\text{--}2 \text{ cm}$ .

$$\Delta\phi_{\text{DFD}} \approx 2 \times 10^{-11} \text{ rad} \times (T/\text{s})^3.$$

*Design B (horizontal Bragg,  $L \sim 1 \text{ m}$ , rotation).* Rotate the bench by  $180^\circ$  about  $\hat{z}$  to flip  $\mathbf{g} \cdot \hat{\mathbf{n}}$ . DFD flips sign; many laser/system alignment systematics do not.

*Design C (tabletop source mass).* Dither a 500 kg tungsten stack at  $R \sim 0.25 \text{ m}$ . Search at the dither frequency; scale with  $g_s/c^2$ .

### V. DISCRIMINANTS FROM GR AND SYSTEMATICS CONTROL

Key orthogonal signatures:

1. **Time scaling:** DFD  $\propto T^3$  vs. GR  $\propto T^2$ .
2. **Orientation:** rotation flips DFD (via  $\mathbf{g} \cdot \hat{\mathbf{n}}$ ), many systematics do not.
3.  **$k$ -reversal:** DFD  $\propto k_{\text{eff}}^2$  (even under  $k_{\text{eff}} \rightarrow -k_{\text{eff}}$ ); laser-phase systematics change sign and cancel.
4. **Recoil dependence:** DFD  $\propto v_r$ ; separate from gravity-gradient terms using velocity selection.
5. **Dual-species:** residual  $\propto (1/m_1 - 1/m_2)$  or the full  $k_{\text{eff}}^2/m$  contrast in Eq. (10); GR null after common-mode rejection.

*Systematics evidence and controls.* Gravity-gradient noise (GGN) from atmosphere and seismic fields sets the long-baseline floor; recent characterizations provide

TABLE I. Systematics overview and kill-switches. The DFD signal alone shows  $T^3$  scaling, rotation sign flip, and even parity under  $k$ -reversal ( $\propto k_{\text{eff}}^2$ ).

Effect	$T$ -scaling	Rotation flip	$k$ -reversal parity
DFD (target)	$T^3$	Yes	Even ( $k_{\text{eff}}^2$ )
Gravity gradient $\Gamma$	$T^2/T^3$ mix	Often No	Mixed
Wavefront curvature / tilt	$T^2$	No	Odd (cancels)
Vibrations (residual)	$\approx T^2$	No	Odd/Even mix
AC Stark / Zeeman	pulse-bounded	No	Design-dependent
Laser phase (uncorrelated)	$T^2$	No	Odd (cancels)

high/low-noise models and motivate underground siting or subtraction.[20, 21] Wavefront aberrations are a leading accuracy term; dedicated measurements and in-situ phase-retrieval methods demonstrate  $< 3 \times 10^{-10} g$  equivalent bias and routes to further reduction.[18, 19] Active isolation routinely delivers  $10^2$ – $10^3$  vertical attenuation at 30 mHz–10 Hz in fieldable systems.[14] Frequency-dependent electronics/Raman-chirp phases are odd-in- $k_{\text{eff}}$  and cancel under  $k$ -reversal with residuals characterized and mitigated.[17, 24] Rotation platforms and mirror-tilt compensation explicitly separate Coriolis/Sagnac terms and have been demonstrated across wide orientation/rotation ranges.[15, 23] Source-mass gravity signals in horizontal/baseline geometries establish lock-in protocols directly applicable to our  $T^3$  search.[22]

## VI. SENSITIVITY SNAPSHOT AND FEASIBILITY

Long-baseline results demonstrate the needed stability and controls: the Stanford 10m fountain achieved long-time point-source interferometry with single-shot acceleration sensitivity at the  $\text{few} \times 10^{-9} g$  level and 1.4 cm arm separation,[7, 8] while dual-species EP tests reached  $\eta \sim 10^{-12}$  with  $2T = 2$  s free fall.[9] VLBAI (Hannover) reports high-flux Rb/Yb sources, 10 m magnetic shielding, and seismic attenuation tailored for long baseline.[10, 11] SYRTE’s absolute gravimeters and mobile surveys document  $\mu\text{Gal}$ -class stability with active vibration isolation.[12–14] These capabilities jointly bound key systematics (vibration, wavefronts, gradients) at or below our target  $|\Delta\phi_{\text{DFD}}| \sim 2 \times 10^{-11}$  rad for  $T \sim 1$  s, and several groups already deploy rotation control and  $k$ -reversal protocols routinely.[15–17]

## VII. DISCUSSION AND OUTLOOK

This work closes the *matter-sector* gap in the DFD experimental program. Together with the cavity–atom redshift comparison (photon sector), matter-wave tests over-constrain the sector coefficients. A null result at or

below the  $|\Delta\Phi|/c^2$  lever arm (after the stated controls) would falsify this DFD sector. Positive detection would present a geometry-locked, scaling-locked deviation from GR that cannot be attributed to standard systematics.

## ACKNOWLEDGMENTS

I thank colleagues in precision atom interferometry for advice on rotation tests, dual-species protocols, and source-mass lock-in strategies.

### Appendix A: Sketch of the $T^3$ derivation from the gradient operator

Write the branch centers as  $z_{A,B}(t) = z_0(t) \pm \frac{1}{2}\Delta z(t)$  with  $\Delta z(t) = v_r t$  for  $0 < t < T$  and  $\Delta z(t) = v_r(2T - t)$  for  $T < t < 2T$ . Expand the field along the arms:

$$\nabla\psi(z_{A,B}) \approx \nabla\psi(z_0) \pm \frac{1}{2}\Delta z \partial_z(\nabla\psi)|_{z_0}. \quad (\text{A1})$$

The constant part  $\nabla\psi(z_0)$  cancels in (4) because  $\int_0^{2T} \Delta p_z dt = 0$  for the piecewise  $\pm\hbar k_{\text{eff}}$  profile. The linear term gives (using Earth field  $\partial_z(\nabla\psi) = -2\Gamma \hat{z}/c^2$  and the kinematic separation implicit in  $\Delta z$ )

$$\begin{aligned} \Delta\phi_{\nabla\psi} &= -\frac{1}{2m} \int dt \left[ \frac{1}{2}\Delta z(t) \partial_z(\nabla\psi) \right] \cdot \Delta\mathbf{p}(t) \\ &\rightarrow \frac{g}{c^2} \frac{\hbar k_{\text{eff}}}{m} \int_0^T t dt + \frac{g}{c^2} \frac{\hbar k_{\text{eff}}}{m} \int_T^{2T} (2T - t) dt \\ &= \frac{\hbar k_{\text{eff}}}{m} \frac{g}{c^2} \left( \frac{T^2}{2} + \frac{T^2}{2} \right) T = \frac{\hbar k_{\text{eff}}}{m} \frac{g}{c^2} T^3, \end{aligned} \quad (\text{A2})$$

and multiplying by the impulsive momentum separation  $\hbar k_{\text{eff}}$  from the light pulses yields (5). A full WKB treatment gives the same result and shows cancellation of the companion  $\psi\nabla^2$  piece for these geometries.

### Appendix B: Figure templates (TikZ/PGFPlots)

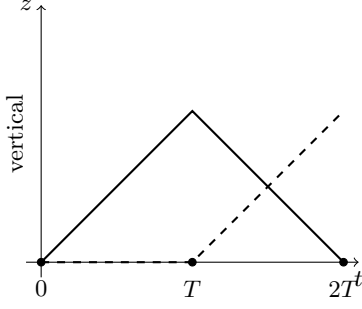


FIG. 1. Light-pulse Mach-Zehnder (Kasevich-Chu) geometry. Solid/dashed are the two arms; pulses at  $0, T, 2T$ .

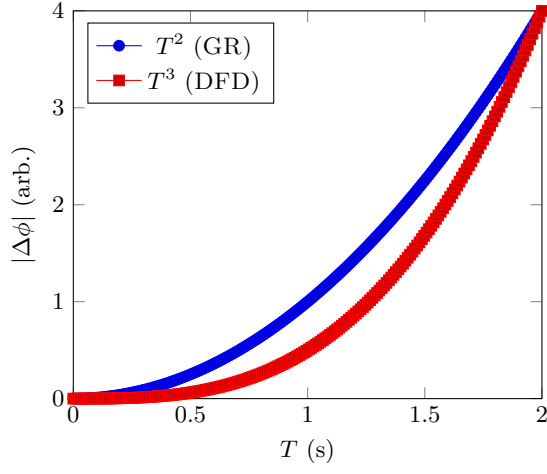


FIG. 2. Scaling discriminator: DFD  $T^3$  vs. GR  $T^2$ .

- 
- [1] M. Kasevich and S. Chu, “Measurement of the gravitational acceleration of an atom with a light-pulse interferometer,” *Appl. Phys. B* **54**, 321 (1992).
  - [2] A. Peters, K. Y. Chung, and S. Chu, “High-precision gravity measurements using atom interferometry,” *Metrologia* **38**, 25 (2001).
  - [3] S. Dimopoulos, P. W. Graham, J. M. Hogan, and M. A. Kasevich, “Atomic gravitational wave interferometric sensor,” *Phys. Rev. D* **78**, 122002 (2008).
  - [4] G. M. Tino and M. A. Kasevich (eds.), *Atom Interferometry* (IOS Press, 2014).
  - [5] A. D. Cronin, J. Schmiedmayer, and D. E. Pritchard, “Optics and interferometry with atoms and molecules,” *Rev. Mod. Phys.* **81**, 1051 (2009).
  - [6] J. M. Hogan, D. M. S. Johnson, and M. A. Kasevich, “Atom interferometry,” *Nat. Phys.* **16**, 913 (2020).
  - [7] S. M. Dickerson *et al.*, “Multiaxis Inertial Sensing with Long-Time Point Source Atom Interferometry,” *Phys. Rev. Lett.* **111**, 083001 (2013).
  - [8] A. Sugarbaker *et al.*, “Enhanced Atom Interferometer Readout through the Application of Phase Shear,” *Phys. Rev. Lett.* **111**, 113002 (2013).
  - [9] P. Asenbaum *et al.*, “Atom-Interferometric Test of the Equivalence Principle at the  $10^{-12}$  Level,” *Phys. Rev. Lett.* **125**, 191101 (2020).
  - [10] D. Schlippert *et al.*, “Very long baseline atom interferometry,” *Proc. SPIE* (2024).
  - [11] D. Schlippert *et al.*, “The Hannover Very Long Baseline Atom Interferometer,” *APS DMP* (2022).
  - [12] P. Gillot *et al.*, “The LNE-SYRTE cold atom gravimeter,” LNE-SYRTE report (2015).
  - [13] X. Wu *et al.*, “Gravity surveys using a mobile atom interferometer,” *Sci. Adv.* **5**, eaax0800 (2019).
  - [14] F. E. Oon *et al.*, “Compact active vibration isolation and tilt stabilization for a transportable quantum gravimeter,” *Phys. Rev. Applied* **18**, 044037 (2022).
  - [15] Q. d’Armagnac de Castanet *et al.*, “Atom interferometry at arbitrary orientations and rotation rates,” *Nat. Commun.* **15**, 6080 (2024).
  - [16] D. Yankelev *et al.*, “Atom interferometry with thousand-fold increase in dynamic range,” *PNAS* **117**, 23414 (2020).
  - [17] B. Cheng *et al.*, “Influence of chirping the Raman lasers in an atom gravimeter,” *Phys. Rev. A* **92**, 063617 (2015).
  - [18] V. Schkolnik *et al.*, “The effect of wavefront aberrations in atom interferometry,” *Appl. Phys. B* **120**, 311 (2015).
  - [19] W. J. Xu *et al.*, “In situ measurement of the wavefront phase shift in an atom interferometer,” *Phys. Rev. Applied* **22**, 054014 (2024).

- [20] J. Carlton *et al.*, “Characterizing atmospheric gravity gradient noise for vertical atom interferometers,” *Phys. Rev. D* **111**, 082003 (2025).
- [21] J. Carlton *et al.*, “Clear skies ahead: atmospheric gravity gradient noise for vertical atom interferometers,” arXiv:2412.05379 (2024).
- [22] G. W. Biedermann *et al.*, “Testing gravity with a horizontal gravity-gradiometer atom interferometer,” *Phys. Rev. A* **91**, 033629 (2015).
- [23] Q. Beaufils *et al.*, “Rotation-related systematic effects in a cold atom accelerometer on a satellite,” *NPJ Microgravity* **9**, 37 (2023).
- [24] Y. Xu *et al.*, “Evaluation of a frequency-dependent phase shift in chirped-Raman atom gravimeters,” *Phys. Rev. A* **110**, 062816 (2024).
- [25] D. Schlippert *et al.*, “Quantum Test of the Universality of Free Fall Using Rubidium and Potassium,” *Phys. Rev. Lett.* **112**, 203002 (2014).
- [26] C. Overstreet *et al.*, “Observation of effective field theory effects in atom interferometry,” *Science* **375**, 226 (2022).
- [27] G. D’Amico, G. Rosi, S. Zhan, L. Cacciapuoti, M. Fattori, and G. M. Tino, “Canceling the Gravity Gradient Phase Shift in Atom Interferometry,” *Phys. Rev. Lett.* **119**, 253201 (2017).
- [28] C. Overstreet, P. Asenbaum, T. Kovachy, R. Notermans, J. M. Hogan, and M. A. Kasevich, “Effective Inertial Frame in an Atom Interferometric Test of the Equivalence Principle,” *Phys. Rev. Lett.* **120**, 183604 (2018).
- [29] A. Roura, “Circumventing Heisenberg’s Uncertainty Principle in Atom Interferometry Tests of the Equivalence Principle,” *Phys. Rev. Lett.* **118**, 160401 (2017).
- [30] P. Asenbaum, C. Overstreet, T. Kovachy, D. D. Brown, J. M. Hogan, and M. A. Kasevich, “Phase Shift in an Atom Interferometer due to Spacetime Curvature across its Wave Function,” *Phys. Rev. Lett.* **118**, 183602 (2017).